

Scaling of Askaryan Pulses @ ARENA

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Topics

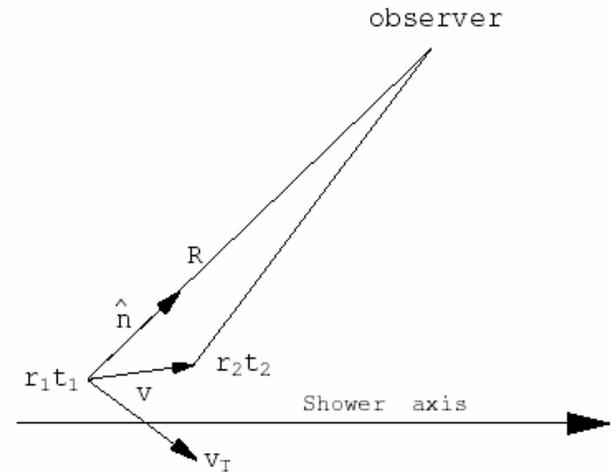
- ReviewRFpulsesfromshowers
- Scalingofuniversalshape
- ApplicationtoAVZ
- Hadronicshowers
- Differentmedia
- Phases&pulseshapes
- Otherissues
- Summary

Radiation from single particles

$$R\vec{E}_\omega = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} R\vec{E}(t)e^{i\omega t} dt ; \quad \text{ZHS:2}$$

$$R\vec{E}(t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} R\vec{E}_\omega e^{-i\omega t} d\omega . \quad \text{0.5}$$

$$\vec{E}(\vec{x}, t) = \frac{q}{c} \left[\frac{\hat{n} \times \{(\hat{n} - \vec{\beta}) \times \dot{\vec{\beta}}\}}{(1 - \vec{\beta} \cdot \hat{n})^3 R} \right]_{\text{ret}}$$



$$R\vec{E}_\omega(\vec{x}) = \frac{1}{\sqrt{2\pi}} \left(\frac{q}{c} \right) e^{ikR} e^{i\omega(t_1 - \hat{n} \cdot \vec{r}_1/c)} \vec{\beta}_\perp \frac{(e^{i\omega\delta t(1 - \hat{n} \cdot \vec{\beta})} - 1)}{1 - \hat{n} \cdot \vec{\beta}}$$

Sumoverparticles/Integrateovershower

$$\hat{n} \times (\hat{n} \times \vec{\beta}) = -\vec{\beta}_\perp$$

$$R\vec{E}_\omega(\vec{x}) = \frac{1}{\sqrt{2\pi}} \left(\frac{q}{c}\right) e^{ikR} e^{i\omega(t_1 - \hat{n} \cdot \vec{r}_1/c)} \vec{\beta}_\perp \frac{(e^{i\omega\delta t(1 - \hat{n} \cdot \vec{\beta})} - 1)}{1 - \hat{n} \cdot \vec{\beta}}$$

$$\vec{E} = \sum_i \vec{E}_i \Rightarrow i\nu \text{Sin}[\theta] \int d^3x dt \rho(\vec{x}, t) e^{i\phi(\theta, \vec{x})}$$

ShowerFront+LightCone

$$E(\theta, \nu) = i\nu \text{Sin}[\theta] \int dy dz \rho(y, z) e^{i\phi(\nu, \theta, y, z)}$$

DensityofExcess
ProjectedTrackLength

Separation of Variables & Form Factor

$$E(\theta, \nu) = i \nu \text{Sin}[\theta] \int d y d z \rho(y, z) e^{i \phi(\nu, \theta, y, z)}$$

$$\rho(y, z) = f_0 f_y(y) f_z(z)$$

$$\phi(\theta, y, z) = \frac{2 \pi \nu}{c} \left(z \left(1 - \frac{\cos \theta}{\beta} \right) + y \frac{\sin \theta}{\beta} \right) = \phi_z + \phi_y$$

$$E(\theta, \nu) = i \nu f_0 \text{Sin}(\theta_c) G_y(\theta, \nu) G_z(\theta, \nu)$$

$$G_y(\theta, \nu) = \int d y f_y(y) e^{i \phi_y(\nu, \theta, y, z)}$$

$$G_z(\theta, \nu) = \int d z f_z(z) e^{i \phi_z(\nu, \theta, y, z)}$$

Phase Factors and Rescaling FF

$$\phi_z = \frac{2\pi\nu}{c} z (\sin \delta \tan \theta_c + (1 - \cos \delta)) \rightarrow \frac{2\pi\nu}{c} z \delta \tan \theta_c$$
$$\phi_y = \frac{2\pi\nu}{c} y (\tan \theta_c \cos \delta + \sin \delta) \rightarrow \frac{2\pi\nu}{c} y \tan \theta_c$$

Most MC's use this

$$\begin{aligned} G_y(\theta, \nu) &= \int dy f_y(y) e^{i\phi_y(\nu, \theta, y, z)} \\ &= \int dy f_y(y) e^{ia_y(\theta) \nu y}; \quad a_y(\theta) = \frac{2\pi}{c} (\tan \theta_c \cos \delta + \sin \delta) \\ &= G_y(0, a_y \nu) \equiv G_y(a_y \nu) \end{aligned}$$

$$G_z(\theta, \nu) = G_z(a_z \nu); \quad a_z(\theta) = \frac{2\pi}{c} (\sin \delta \tan \theta_c + (1 - \cos \delta))$$

Normalization and Determination of G_y, G_z

$$E(\theta, \nu) = i\nu f_0 \sin(\theta_c) G_y(a_y \nu) G_z(a_z \nu)$$

$$G_y(0) = 1, \quad G_z(0) = 1$$

$$a_z(\delta = 0) = 0$$

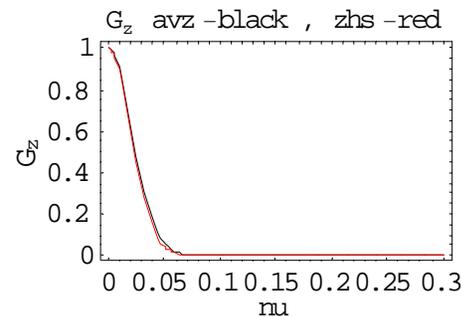
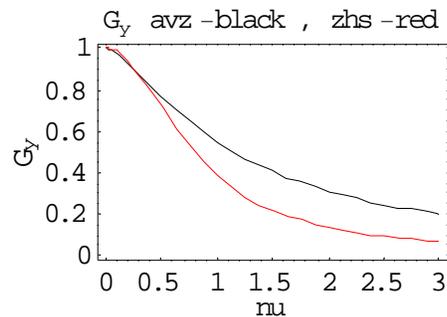
$$\begin{aligned} f_0 &= \lim_{\nu \rightarrow 0} \frac{|E(0, \nu)|}{\nu \sin(\theta_c)} \\ G_y(\nu) &= \frac{|E(0, \nu)|}{\nu f_0 \sin(\theta_c)} \\ \left(G_z(a_z \nu) = \frac{-iE(\theta, \nu)}{\nu f_0 \sin(\theta) G_y(a_y \nu)} \right) \Big|_{\theta} \end{aligned}$$

AVZ,ZHS,(G4later)

$$E(\nu, 0, \text{ice}) = 2.53 \times 10^{-7} \frac{\nu}{\nu_0} \frac{1}{1. + \left(\frac{\nu}{\nu_0}\right)^{1.44}},$$

$$G_y(\nu) = \frac{1}{1. + \left(\frac{\nu}{\nu_0}\right)^{1.44}}, \quad f_0 = \frac{2.53 \times 10^{-7}}{\nu_0 \text{Sin}[\theta_{c0}]} \text{ (units?)}.$$

AVZ Use a Gaussian, and give the half-width of the Cerenkov cone
 at 1GHz as 3.72° for a 1PeV (3.1m) shower.
 Determine G_z as FT of gaussian



Scaling vs AVZ

Field calculation is integral over shower profile

$$E = v \sin[\theta] \int dy dz \rho(y, z) e^{i\phi(\theta, y, z)}$$

Separation of shower profile

$$\rho(y, z) = f_0 f_y(y) f_z(z)$$

Separation of phase factors

$$\phi(\theta, y, z) = \frac{2\pi v}{c} \left(z \left(1 - \frac{\cos \theta}{\beta} \right) + y \frac{\sin \theta}{\beta} \right) = \phi_z + \phi_y$$

$$\phi_z = \frac{2\pi v}{c} z (\sin \delta \tan \theta_c + (1 - \cos \delta))$$

$$\phi_y = \frac{2\pi v}{c} y (\tan \theta_c \cos \delta + \sin \delta)$$

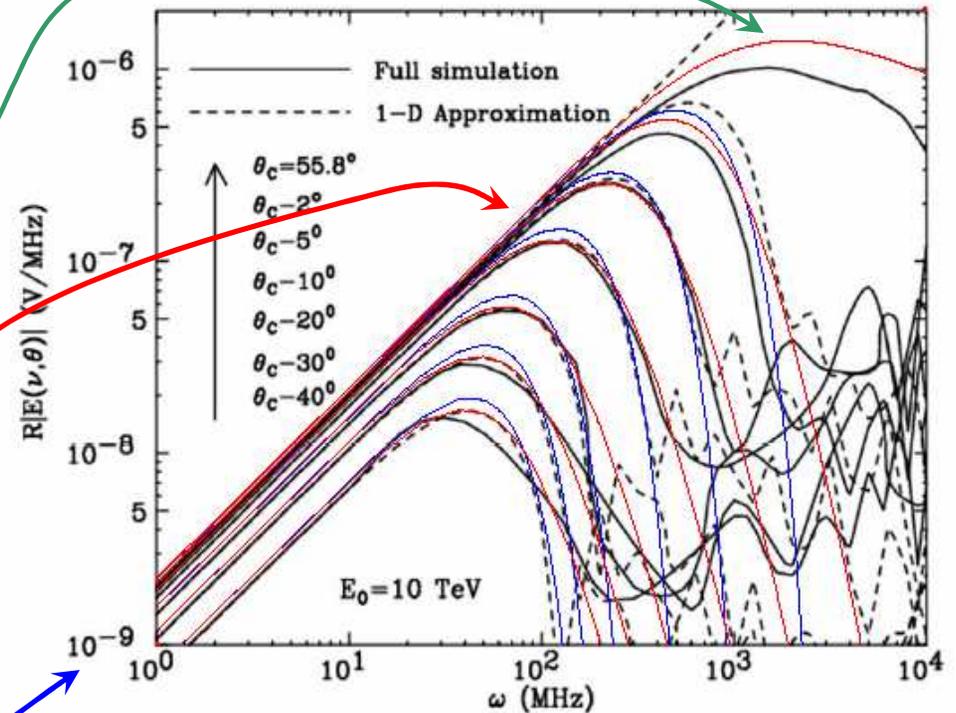
Separation of form factors

$$E(v, \delta) = f_0 v \sin[\theta] G_y(v_y) G_z(v_z)$$

With scaled frequencies

$$v_y = v \left(\cos \delta + \frac{\sin \delta}{\tan \theta_c} \right) \approx v \left(1 + \frac{\delta}{\tan \theta_c} \right)$$

$$v_z = v l \left(\sin \delta + \frac{(1 - \cos \delta)}{\tan \theta_c} \right) \approx v l \delta$$



- Adapted from Alvarez, Vazquez, Zas
- "Fullsim" is approx
- Blue – Gaussian for $f_z(z)$, AVZ approx for G
- Red – Griessen for $f_z(z)$

Highfrequency

- Large phases from individual particles... slow to average
- Initial particle track may belong, and be noticed for $E_s < 1 \text{ TeV}$
- Not clear to me that AVZ form is not contaminated (form was derived at 10 TeV)

Phases

- later

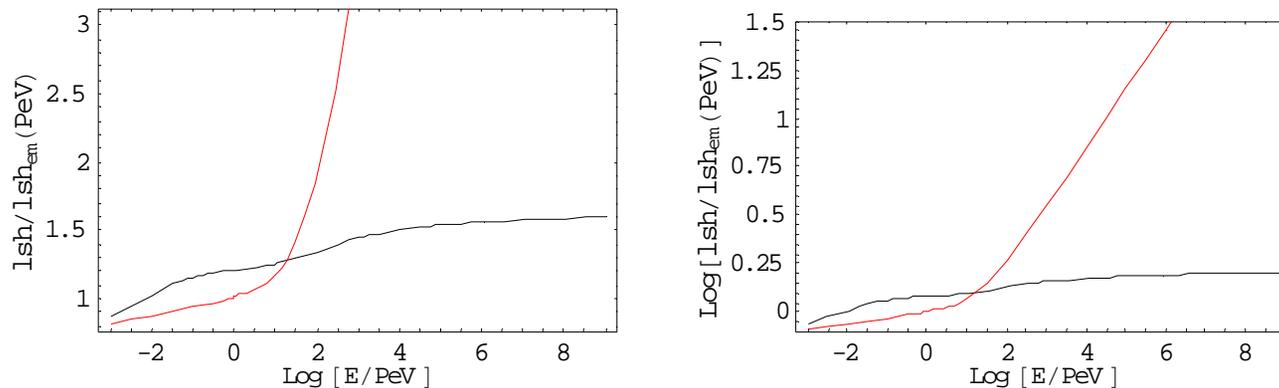
Scaling with energy and type. (AZ)

$$f_0: \quad l_{\text{tot}} = \frac{E}{dE/dx}$$

G_y : R_m doesn't change

G_z : $l_{\text{sh}}(i, E)$

Comparison of em and hadronic shower length



1. LPM effect (Note: limited at $\sim 10^{20}$ eV by $\gamma\gamma$ interactions (BR, K))
2. At PeV proton showers are longer
3. Hadronic shower size asymptotes at about 93% of e⁺-shower

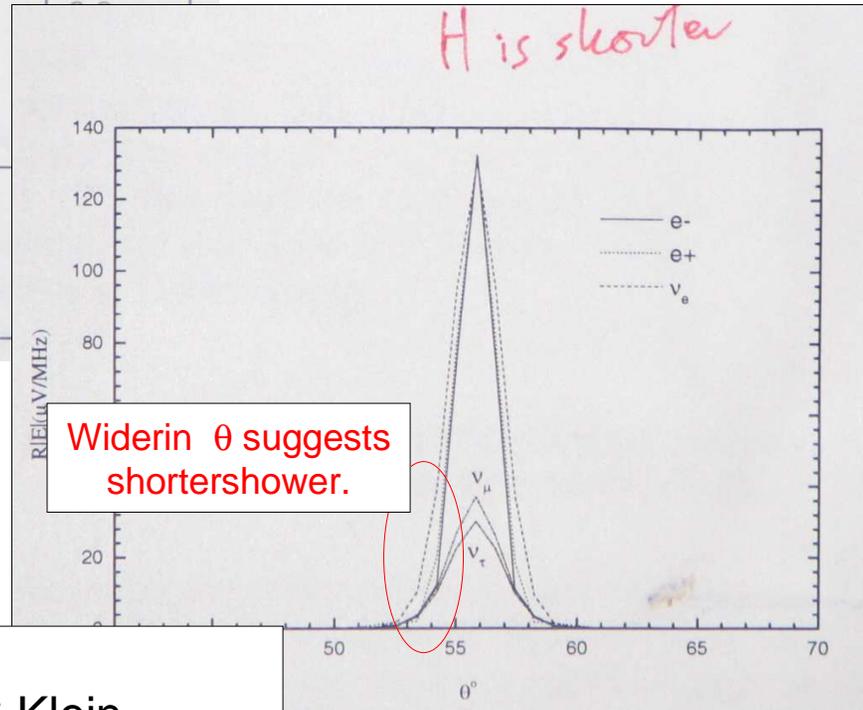
Hadronic showersII.(HM)

E	name	Δr_z	$R E 1$	w	$b(w) * 1e5$	$R E 2$
10^2	e^-	$1.33E5$	13	0.90	9.9	13
90	π^-	$1.04E5$	9.2	0.85	8.8	9.1
10^2	π^+	$1.19E5$	11	0.86	9.0	11
90	p	90007	6.9	0.80	7.5	
10^2	ν_e	$1.26E5$	12	0.87	9.3	
10^2	ν_μ	39982	3.5	0.84	8.5	
10^2	ν_τ	36800	3.0	0.82	8.0	
10^3	e^-	$1.33E6$	132	0.90	9.9	
	ν_e	$1.32E6$	131	0.89	9.7	
	ν_μ	$4.08E5$	37	0.86	9.0	
	ν_τ	$3.67E5$	30	0.82	8.1	

Excesscharge

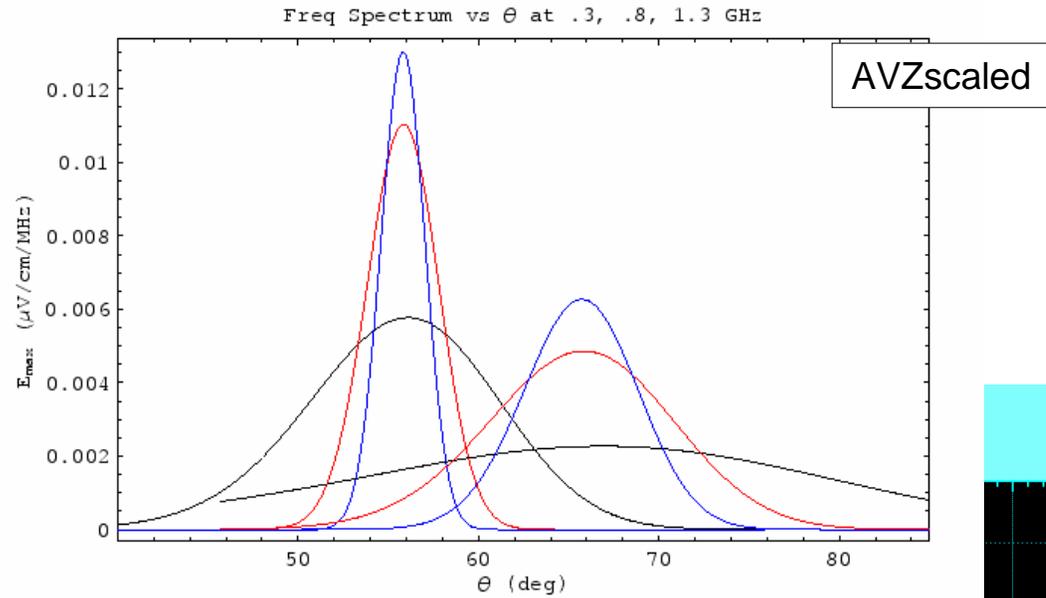
E-field@1GHz

Containedcharge@1/GHz



- Sizesimilar to AZ
- Same R_M ? (Affects G_y) – also see S. Klein
- Difference between proton and jet (G_z)?
- What about charm (~10%)?

Scalingwithmedium

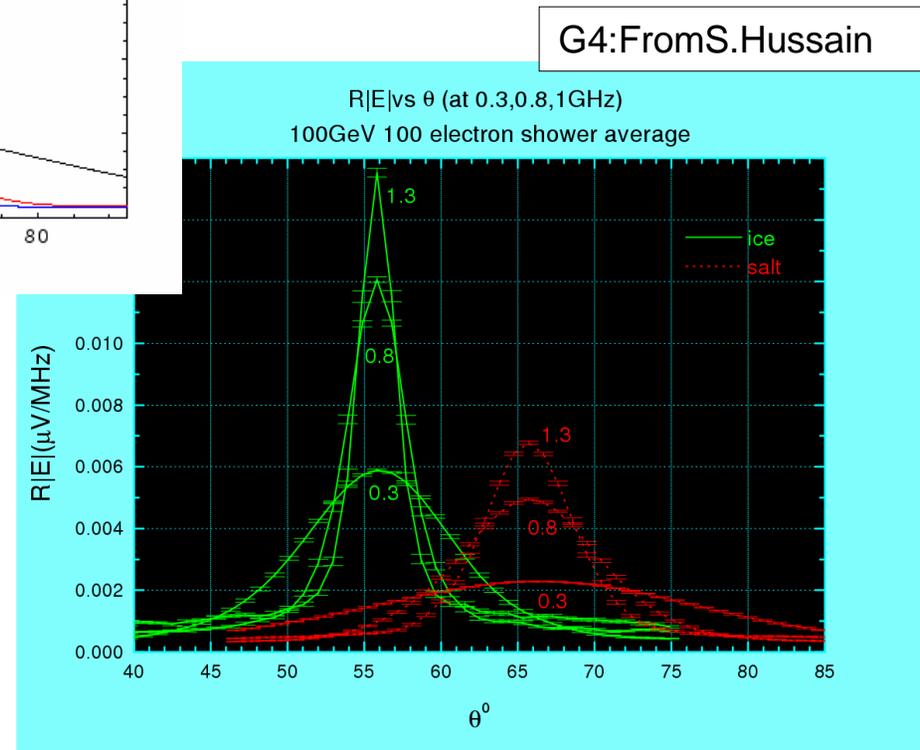


Yieldofprojectedtracklength

$$f_0: l_{\text{tot}} = \frac{\alpha_i E}{dE/dx}$$

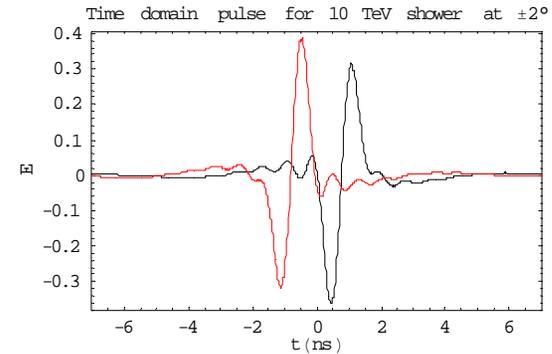
$$G_y: R_M \sim \frac{E_s}{\rho dE/dx}$$

$$G_z: l_{\text{sh}} \sim x \sim \frac{X}{\rho}$$

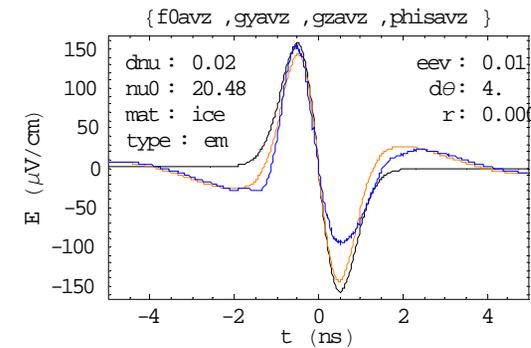


Phases & Pulse shapes

Hussain: G4 10 TeV shower, reconstructed complex
 $E(\omega)$ at $\theta = (-4, -2, 0, 2, 4)$, FFT $\Rightarrow E(t)$



Comparison to AVZ scaled at $\theta = 4$. AVZ pulse shape
 symmetric bipolar. Comparison shown with and
 Without filter



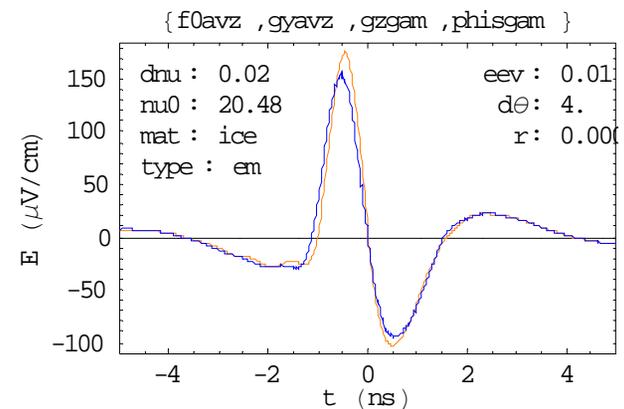
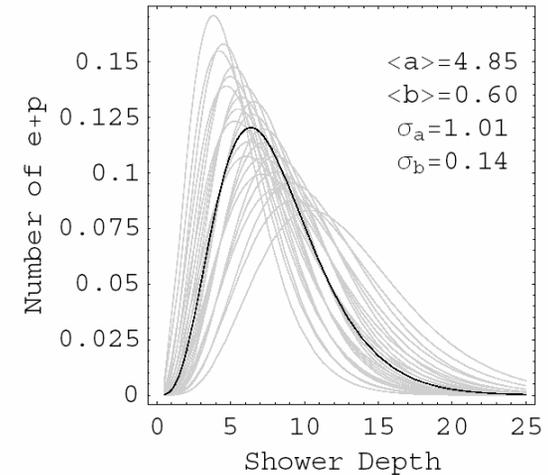
If $\rho(t)$ symmetric, G_z is real, $\phi = 0$. $i \nu G_z$ will transform to
 time derivative of ρ will be bipolar symmetric. If not,
 then $\phi \neq 0$, asymmetric shape. Change sign with θ .

Gammaisabeterpulseshape

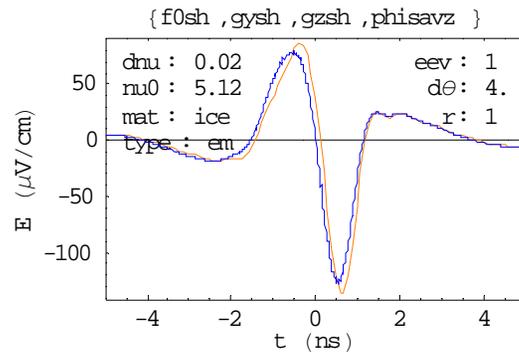
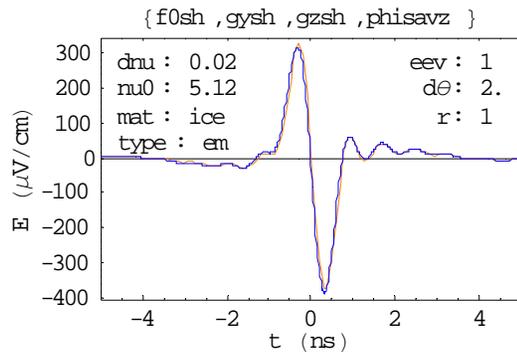
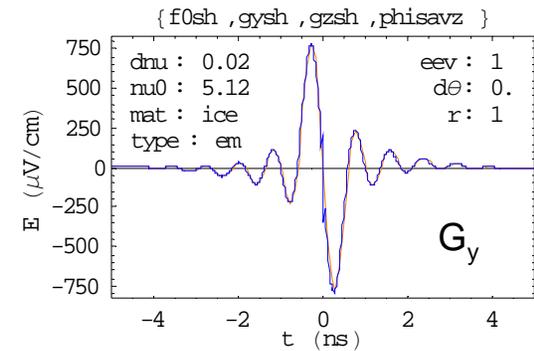
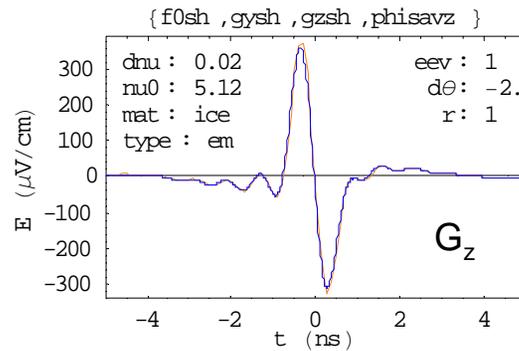
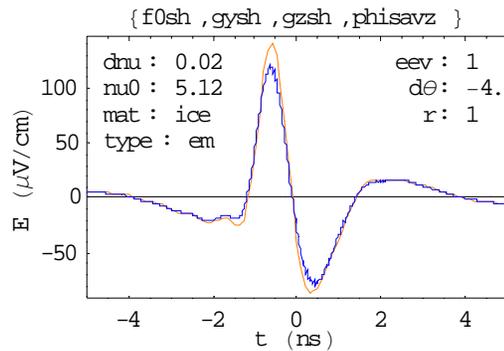
- Without phases AVZ effectively gaussian shower profile.
- Gamma function is approximation to shower profile. $\rho_z(t) = N t^{a-1} e^{-bt}$
- Causal, analytic phase, easy to scale.

$$\begin{aligned}
 g(\omega) &= \int_0^\infty dt e^{i\omega t} N t^{a-1} e^{-bt} \\
 &= \left(\frac{1}{1-i\frac{\omega}{b}} \right)^a \\
 &= \left(\frac{b}{\sqrt{b^2+\omega^2}} \right)^a e^{ia \text{ArcTan}(\frac{\omega}{b})}
 \end{aligned}$$

- Comparison to G4



Determine G_y, G_z from G4shower&test



Spectrum at $\theta = (-4, -2, 0, 2, 4)$ from Hussein.

Use $\theta = 0$ to get G_y

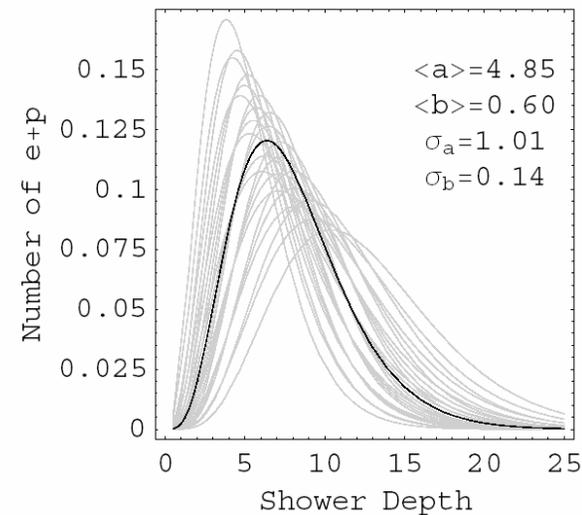
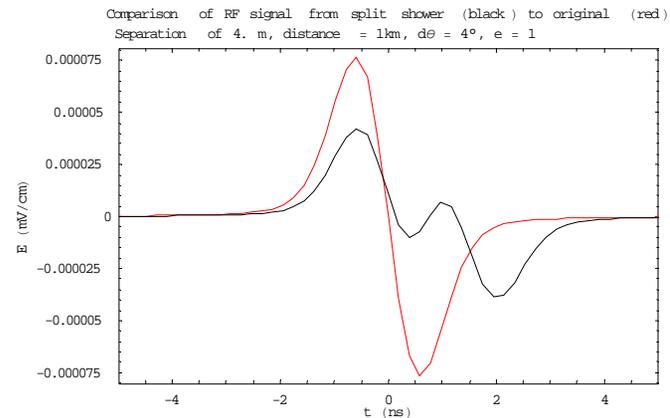
Use $\theta = -2$ to get G_z

$\theta = (-4, 2, 4)$ from scaling, with phase part of G

z

Multipleshowers&fluctuations

- ν_e CCevents
- Multipleinteractions
- Modelfluctuationsby addingsubshowers frombeginningof shower?
- Fluctuationsfrom distributiona,b?



Acoustic(inprogress)

- Same as radio....except
 - scalar vs vector
 - total track length
 - v/c is smaller
 - Attenuation is different
 - experiment bandwidth

Summary

- Scaling of RF pulse to universal shape describes much previous and ongoing work.
- Hadronic showers need another look.
- Different media can be understood
- Better shower shape & treatment of phases needed
- Multiple showers, ...